## Chemical Evolution of the Universe Problem sheet 9

- 1. Let  $m = \frac{M}{M_{\odot}}$ . Consider the Salpeter Initial Mass Function (IMF), given by  $\Phi(m) = A m^{-2.35}$ , where A is a constant, with mass limits  $m_l = 0.1$  and  $m_u = 100$ .
  - (a) Compute A from the standard normalisation condition  $\int_{m_l}^{m_u} m \Phi(m) dm = 1$ .
  - (b) Calculate the median mass by number,  $m_{N,1/2}$ , defined by  $f_N(m_{N,1/2}) = \frac{1}{2}$ .
  - (c) Calculate the fraction of mass contained in stars with  $m > m_{N,1/2}$ , i.e.  $f_M(m_{N,1/2})$ .
  - (d) Now compute the median mass by mass,  $m_{M,1/2}$ , defined by  $f_M(m_{M,1/2}) = \frac{1}{2}$ .
  - (e) Consider a galaxy with a constant star formation rate of  $\psi = 20 M_{\odot}/\text{yr}$ . Assuming that every star with m > 8 explodes as a supernova (SN), how many years do we have to wait (on average) in order to see one SN in this galaxy?

## 6 points

2. Compute the lock-up and return fractions,  $\alpha$  and R, for Salpeter's IMF (see above) in the instantaneous recycling approximation, assuming m = 1 to be the boundary between stars that have already died and those that have not.

## 2 points

- 3. Consider a stellar cluster of age  $t_C$  and some star formation history  $\psi(t)$ . Let us assume that we know the distance to this cluster and that we have hence been able to measure the present-day mass function of its main sequence stars, N(M). How can these data be used to constrain the unknown IMF of the cluster,  $\Phi(M)$  (assumed to be constant as a function of time)?
  - (a) Assume that  $\tau(M) \gg t_C$  for all  $M < M_1$ , where  $\tau(M)$  is the main sequence lifetime of a star of mass M. What is the relationship between N(M) and  $\Phi(M)$  for  $M < M_1$ ?
  - (b) Further assume that  $\tau(M) \ll t_C$  for  $M > M_2$  and that  $\psi(t)$  is approximately constant over the small time interval  $[t_C \tau(M_2), t_C]$ . What is the relationship between N(M) and  $\Phi(M)$  for  $M > M_2$ ?
  - (c) Further assume a linear relationship between  $\log \Phi$  and  $\log M$  in the intermediate mass regime  $M_1 < M < M_2$ . What is the slope of this relationship?

## 4 points